

# Non-integrability criterion for homogeneous Hamiltonian systems via blowing-up theory of singularities

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# Hamiltonian system and its integrability

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- Hamiltonian system :

$$\frac{dq_j}{dt} = \frac{\partial H}{\partial p_j}(p, q), \quad \frac{dp_j}{dt} = -\frac{\partial H}{\partial q_j}(p, q) \quad (j = 1, \dots, k) \quad (1)$$

where  $p = (p_1, \dots, p_k)$ ,  $q = (q_1, \dots, q_k)$ ,  $H : \mathbb{R}^{2k} \rightarrow \mathbb{R}$ .

- Hamiltonian system (1) is integrable  $\iff$  there are  $k$  first integrals  $F_1 (= H), F_2, \dots, F_k$  such that  $dF_1, \dots, dF_k$  are linearly independent a.e. and that  $\{F_i, F_j\} = 0$  for any  $i, j = 1, \dots, k$ .
- The dynamics of the integrable Hamiltonian systems are well understood because of the Liouville-Arnold theorem.
- The dynamics of the non-integrable Hamiltonian systems may be “chaotic”.
- **Problem: distinguish between integrable and non-integrable Hamiltonian systems.**

## Brief history

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- Bruns (1887) proved  $\nexists$  algebraic first integral in the 3BP.
- Poincaré (around 1890) proved  $\nexists$  analytic first integral in the R3BP.
- Kovalevskaya (1889) discovered an integrable parameter in the rigid body model by focusing on the property of the singularity.
- Ziglin (1982 –) provided a criterion for non-integrability by using Monodromy matrix.
- Yoshida (1986 –) provided a criterion for non-integrability of homogeneous Hamiltonian systems.
- Morales-Ruiz & Ramis (1999 –) extended the Ziglin analysis by using the differential Galois theory.
- Maciejewski (2011) proved meromorphic non-integrability of the P3BP for any masses by applying the Morales-Ramis theory.

# Goal

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- Goal: give a criterion of the non-integrability of the homogeneous Hamiltonian systems with two degrees of freedom from **a new approach**.

## Homogeneous Hamiltonian system

Consider a homogeneous Hamiltonian system with two degrees of freedom:

$$H(\mathbf{p}, \mathbf{q}) = \frac{1}{2} \|\mathbf{p}\|^2 + U(\mathbf{q}) \quad ((\mathbf{p}, \mathbf{q}) \in \mathbb{R}^2 \times \mathbb{R}^2)$$

where  $U$  is a homogeneous potential with degree  $\beta (\in \mathbb{R})$ :

$$U(\lambda \mathbf{q}) = \lambda^\beta U(\mathbf{q}) \quad (\forall \mathbf{q} \in \mathbb{R}^2 \setminus \{\mathbf{0}\}, \forall \lambda > 0).$$

Let  $V(\theta) = U(\cos \theta, \sin \theta)$ .

## Example(The isosceles three-body problem)

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Consider the isosceles three-body problem.

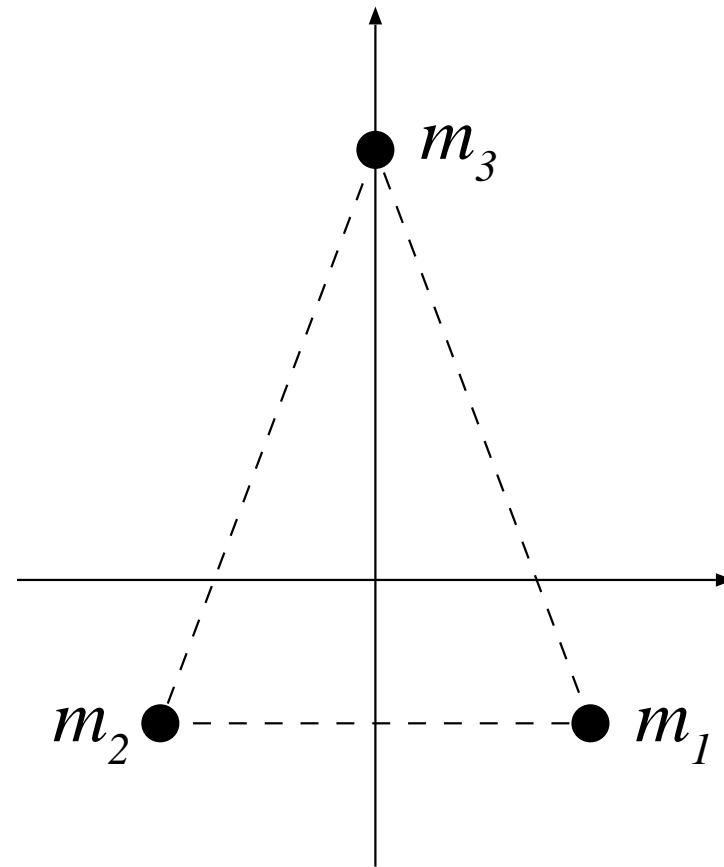
Assume that  $m_1 = m_2, m_3 = \alpha m_1$ .

This model is governed by the homogeneous Hamiltonian system with the potential energy

$$U(\mathbf{q}) = -\frac{1}{q_1} - \frac{4\alpha^{3/2}}{\sqrt{\alpha q_1^2 + (\alpha + 2)q_2^2}}.$$

$$\beta = -1.$$

$$V(\theta) = -\sec \theta - \frac{4\alpha^{3/2}}{\sqrt{\alpha + 2 \sin^2 \theta}}.$$





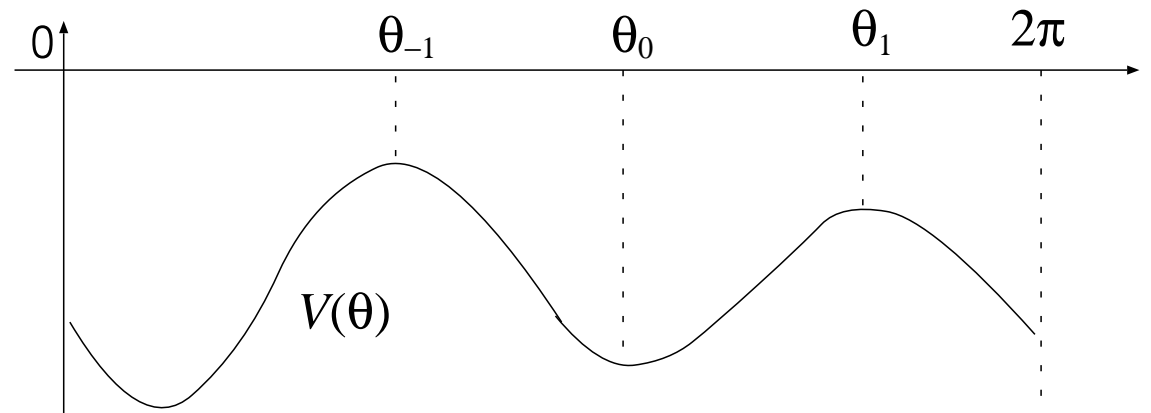
# Main result

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Theorem Assume the following:

1.  $\beta \in \mathbb{R} \setminus \{-2, 0\}$ ;
2.  $\exists \theta_{-1} < \exists \theta_0 < \exists \theta_1$  s.t.  $\frac{\partial V}{\partial \theta}(\theta_l) = 0$ ;
3.  $V(\theta) < 0$  on  $[\theta_{-1}, \theta_1]$ ;
4.  $\frac{\partial V}{\partial \theta}(\theta) \neq 0$  on  $(\theta_{-1}, \theta_0) \cup (\theta_0, \theta_1)$ ;
5.  $\frac{\partial^2 V}{\partial \theta^2}(\theta_{\pm 1}) < 0$ ;
6.  $-\frac{1}{8}(\beta + 2)^2 V(\theta_0) < \frac{\partial^2 V}{\partial \theta^2}(\theta_0)$ .

Then the homogeneous Hamiltonian system has no meromorphic first integral independent from  $H$ .



## Remark

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In the case of  $\beta = -2$ , the Hamiltonian system is always integrable. Because a function

$$G(\mathbf{p}, \mathbf{q}) = (\mathbf{q} \cdot \mathbf{p})^2 - 2\|\mathbf{q}\|^2 H(\mathbf{p}, \mathbf{q})$$

is a first integral independent from  $H$ .

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# McGehee coordinates

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We mainly consider the case of  $\beta < 0$ .

McGehee coordinates:  $(r, \theta, v, w)$  and  $\tau$

$$\mathbf{q} = r(\cos \theta, \sin \theta),$$

$$\mathbf{p} = r^{\beta/2}(v(\cos \theta, \sin \theta) + w(-\sin \theta, \cos \theta))$$

$$dt = r^{1-\beta/2}d\tau.$$

Then the canonical equations become

$$\frac{dr}{d\tau} = rv \tag{2}$$

$$\frac{d\theta}{d\tau} = w \tag{3}$$

$$\frac{dv}{d\tau} = -\frac{\beta}{2}v^2 + w^2 - \beta V(\theta) \tag{4}$$

$$\frac{dw}{d\tau} = -\left(\frac{\beta}{2} + 1\right)vw - \frac{\partial V}{\partial \theta}(\theta) \tag{5}$$

$\mathbf{q} = 0$  is singularity but  $r = 0$  is not singular in these differential equations (2)-(5).

## Energy and Collision manifold

In these coordinates the total energy is

$$h = r^\beta \left( \frac{v^2 + w^2}{2} + V(\theta) \right). \quad (6)$$

We fix  $h \neq 0$  and regard  $r$  as a function of  $(\theta, v, w)$ .

We consider the 3-dimensional dynamics.

The set

$$\mathcal{M} = \left\{ (\theta, v, w) \mid \frac{v^2 + w^2}{2} + V(\theta) = 0 \right\}$$

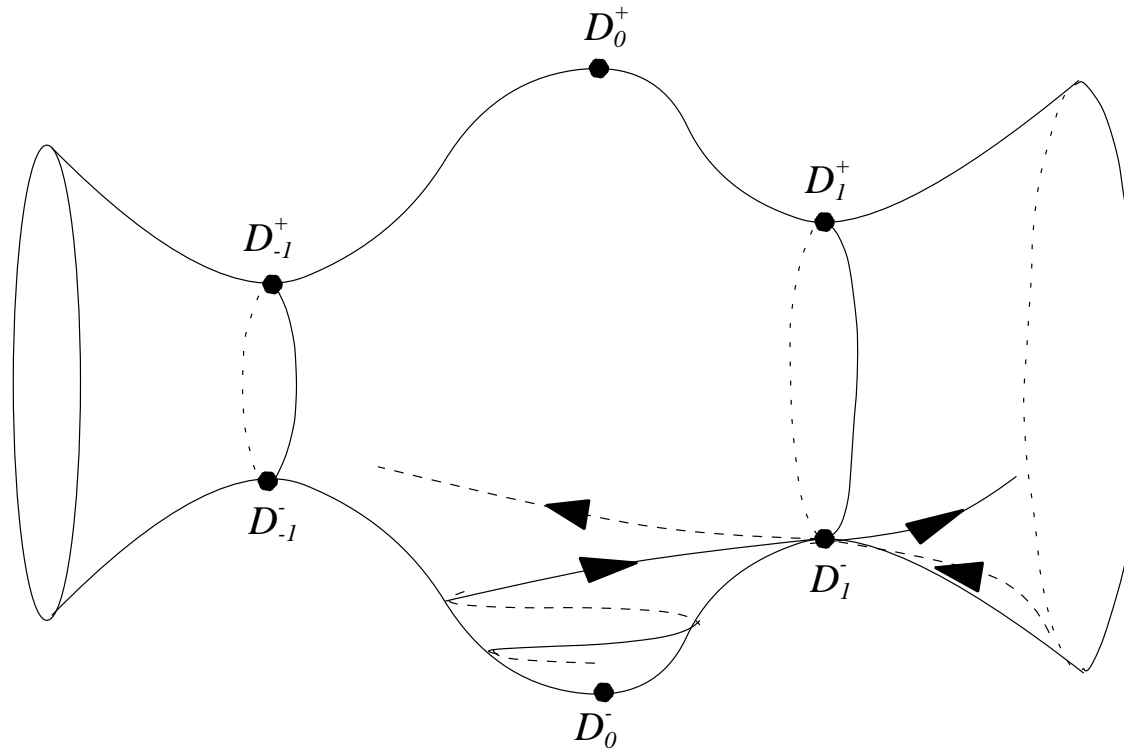
is invariant. In the case of the  $n$ -body problem,  $\mathcal{M}$  is called collision manifold.

Since we fix the energy, as  $r \rightarrow \mathbf{0}$  ( $\mathbf{q} \rightarrow \mathbf{0}$ ), the orbit converges to  $\mathcal{M}$  in the McGehee coordinates.

## Equilibrium points

Recall that  $\theta_l$  are a critical point of  $V$ , i.e.  $\frac{\partial V}{\partial \theta}(\theta_l) = 0$ .

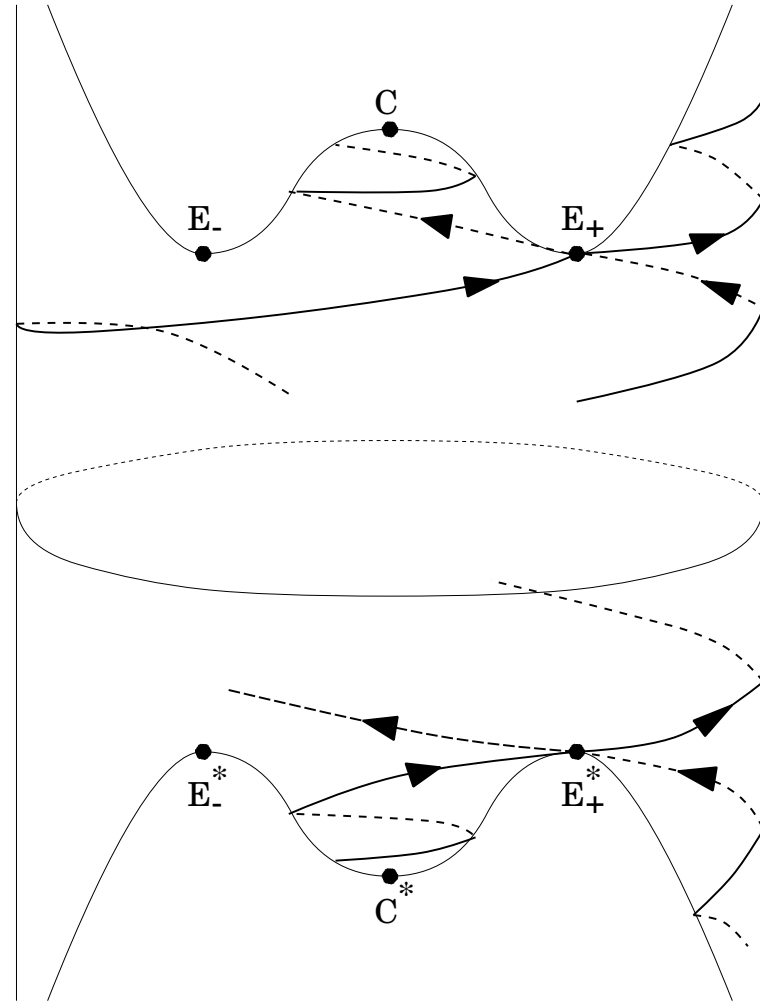
Then  $D_l^\pm = (\theta_l, \pm\sqrt{-2V(\theta_l)}, 0) \in \mathcal{M}$  are equilibrium points.



# The case of the isosceles three-body problem

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The invariant manifold (collision manifold)  $\mathcal{M}$  for the isosceles three-body problem is like this figure:



## Case of $\beta > 0$

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In the case of  $\beta > 0$ , we replace  $r$  with  $R = r^{-1}$ .

The equation  $\frac{dr}{d\tau} = rv$  becomes  $\frac{dR}{d\tau} = -Rv$ .

We can define an invariant manifold corresponding to  $R \rightarrow 0$  and we can discuss a similar argument as the case of  $\beta < 0$ .



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## Proof (homogeneous property)

We give the outline of the proof for  $-2 < \beta < 0$ . The other cases are similar (some signs change in the computation).

Assume that  $\Phi(\mathbf{p}, \mathbf{q})$  is a meromorphic first integral where  $(\mathbf{p}, \mathbf{q})$  are the original coordinates.

From the homogeneous property

(if  $(\mathbf{p}(t), \mathbf{q}(t))$  is a solution, so is  $(c^{\beta/2}\mathbf{p}(c^{\beta/2-2}t), c\mathbf{q}(c^{\beta/2-2}t))$  for any constant  $c > 0$ ),

we can assume that  $\Phi$  satisfies  $\Phi(c^{\beta/2}\mathbf{p}, c\mathbf{q}) = c^\rho\Phi(\mathbf{p}, \mathbf{q})$  without loss of generality.

In the McGehee coordinates, this property corresponds to the fact that  $\Phi$  can be represented as  $\Phi = r^\rho g(\theta, v, w)$ .

## Proof(Coordinates)

We use the coordinates  $(\theta, z, w)$  where  $z = \frac{v^2+w^2}{2} + V(\theta)$ . These are analytic near the equilibrium points. The energy is

$$h = r^\beta z. \quad (7)$$

We consider the Laurent series of  $g$  at  $z = 0$  with respect to  $z$ :

$$g = \sum_{k=\mu}^{\infty} \gamma_k(\theta, w) z^k \quad (\gamma_\mu \neq 0).$$

From (7), we get  $\Phi = \left(\frac{h}{z}\right)^{\frac{\rho}{2\beta}} \sum_{k=\mu}^{\infty} \gamma_k(\theta, w) z^k$ .

The lowest order of  $z$  is  $\mu - \frac{\rho}{2\beta}$ .

## Proof(the case of $\mu - \frac{\rho}{2\beta} < 0$ )

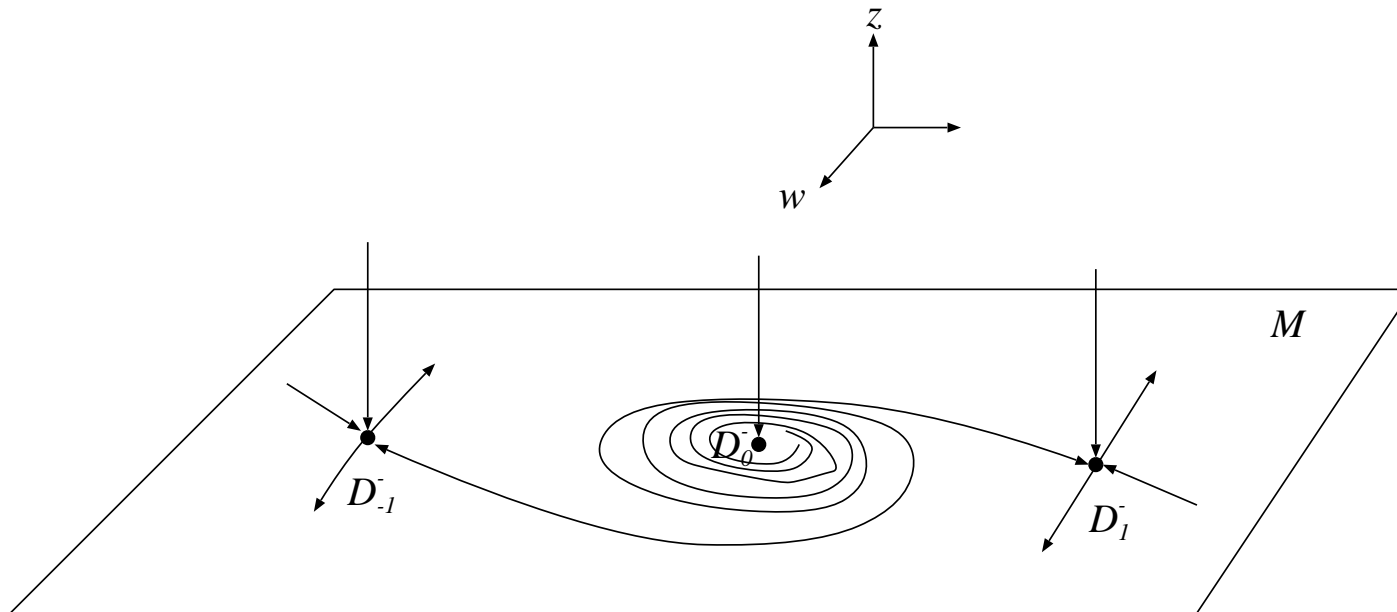
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We first consider the case of  $\mu - \frac{\rho}{2\beta} < 0$ .

Lemma:  $\gamma_\mu$  is zero on  $W^u(D_1^-)$ .

$W^u(D_0^-)$  is an open set of  $\mathcal{M}$ . Hence  $\gamma_\mu \equiv 0$ .

This contradicts the assumption.



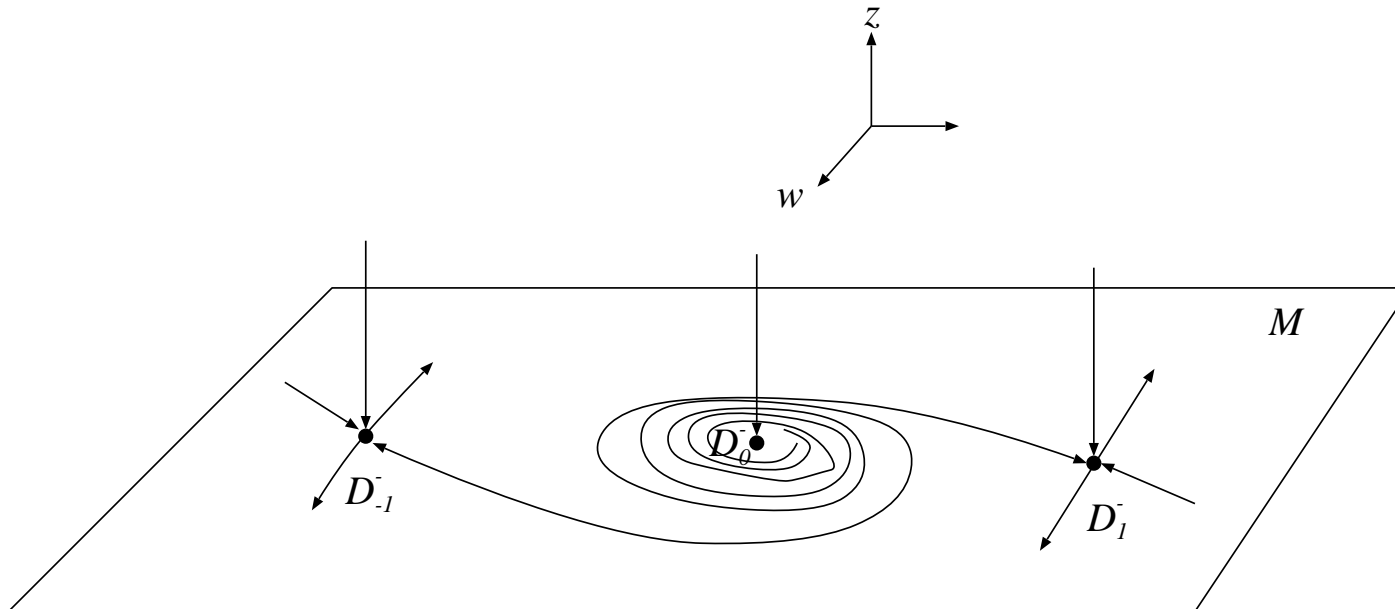
## Proof(the case of $\mu - \frac{\rho}{2\beta} > 0$ )

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We consider the case of  $\mu - \frac{\rho}{2\beta} > 0$ .

Lemma:  $\gamma_\mu$  is zero on  $W^s(D_1^-)$

From assumption 6:  $(-\frac{1}{8}(\beta + 2)^2 V(\theta_0) < \frac{\partial^2 V}{\partial \theta^2}(\theta_0))$ , the dynamics near  $D_0^-$  on  $\mathcal{M}$  is unstable focus.  $W^s(D_1^-)$  is a spiral curve near  $D_0^-$ . Hence  $\gamma_\mu \equiv 0$ .



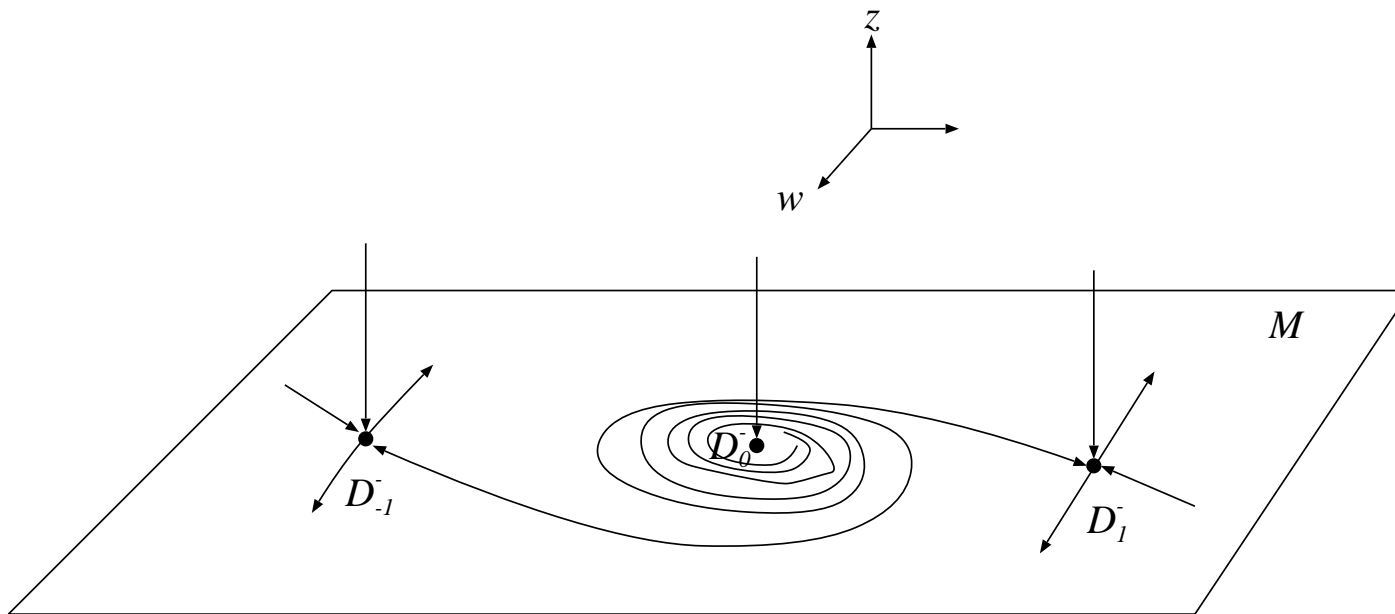
Proof(the case of  $\mu - \frac{\rho}{2\beta} = 0$ )

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In the case of  $\mu - \frac{\rho}{2\beta} = 0$ ,

Lemma:  $\gamma_\mu$  is a constant on  $W^{s/u}(D_l^-)$ .

Therefore  $\gamma_\mu \equiv c$ . If  $\Phi$  is not constant, by considering  $\Phi - c$ , this case can be reduced to the case of  $\mu - \frac{\rho}{2\beta} > 0$ .



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## Non-integrability of the isosceles three-body problem

The function  $V$  of this problem is

$$V(\theta) = -\sec \theta - \frac{4\alpha^{3/2}}{\sqrt{\alpha + 2\sin^2 \theta}}.$$

By applying our theorem, we obtain the following:

### Theorem 2

Assume that  $\alpha < \frac{55}{4}$ . Then the isosceles three-body problem is non-integrable. i.e. there is no meromorphic first integral independent from the energy.



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## Yoshida coefficient

We call a point  $\mathbf{c} \in \mathbb{R}^2$  the Darboux point if  $\nabla U(\mathbf{c}) = \mathbf{c}$ . In the case of  $n$ -body problem,  $\mathbf{c}$  is called a central configuration. The eigenvalues of the Hessian matrix  $D^2U(\mathbf{c})$  at the Darboux point  $\mathbf{c}$  are called the Yoshida coefficients.

Since  $U(\mathbf{c})$  is homogeneous with degree  $\beta$ , one of Yoshida coefficients is  $\beta - 1$ .

The other Yoshida coefficient is

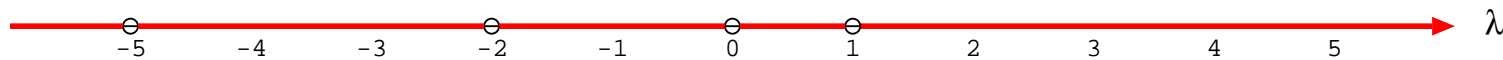
$$\lambda = \beta^{-1} V(\theta_c)^{-1} \frac{\partial^2 V}{\partial \theta^2}(\theta_c) + 1$$

in the polar coordinates where  $\frac{\partial V}{\partial \theta}(\theta_c) = 0$ .

# Yoshida coefficient and integrability

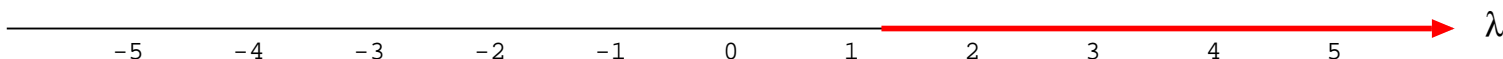
The Morales-Ramis theorem (the differential Galois theory) proves non-integrability if one of the Yoshida coefficient is not in a certain set of rational numbers. For example, in the case of  $\beta = -1$ , according to the Moreles-Ramis theorem, the homogeneous Hamiltonian system is non-integrable if  $\lambda$  is not in

$$\left\{-\frac{1}{2}p(p-3) \mid p \in \mathbb{Z}\right\} = \{1, 0, -2, -5, -9, \dots\}.$$



In our theorem the assumption 6 is

$$-\frac{1}{8}(\beta + 2)^2 > (\lambda - 1)\beta \quad (\lambda > 9/8 \text{ if } \beta = -1).$$



In the case of the isosceles three-body problem,

- Our theorem: non-integrability for  $\alpha < \frac{55}{4}$
- M-R theory: non-integrability for any  $\alpha$ .

## Our theorem v.s. Morales-Ramis theory

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- Our theorem can be applied to  $\beta \in \mathbb{R} \setminus \{-2, 0\}$  while M-R theory can be applied to  $\beta \in \mathbb{Z} \setminus \{-2, 0\}$ .
- In the case of integer  $\beta$ , M-R theory is stronger.
- Our theorem can be applied to **two** degrees of freedom while M-R theory can be applied to **any** degrees of freedom.
- Our function class of first integrals is bigger: we prove the non-existence of first integral which is **meromorphic as a real function**, while M-R theory prove the non-existence of first integrals which is **meromorphic as a complex function**.
- Our proof is **simpler and based on dynamics** (the behavior of stable and unstable manifolds). M-R's method is **far from the theory of the dynamics**.

Thank you for your attention.